3.23 Electrical, Optical, and Magnetic Properties of Materials Fall 2007

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3.23 Fall 2007 – Lecture 4 (LOSE TO (OLLAPSE

The collapse of the wavefunction

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Travel

- Office hour (this time only):
 - This Friday, Sep 21, 4pm
 - (instread of Mon, Sep 24, 4pm)

Last time: Wave mechanics

- 1. The ket $|\Psi
 angle$ describe the system
- 2. The evolution is deterministic, but it applies to stochastic events
- 3. Classical quantities are replaced by operators
- 4. The results of measurements are eigenvalues, and the ket collapses in an eigenvector

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Commuting Hermitian operators have a set of
common eigenfunctions

$$\hat{A}, \hat{B} [\hat{A}, \hat{B}] = 0 \Rightarrow \hat{A}\hat{B} = \hat{B}\hat{A} \hat{A} |\Psi_n \rangle = a_n |\Psi_n\rangle$$

 $\hat{I}) \hat{A}\hat{B} |\Psi_n \rangle = \hat{B}\hat{A} |\Psi_n \rangle = \hat{B} a_n |\Psi_n \rangle = a_n \hat{B} |\Psi_n\rangle$
 $\hat{A} (\hat{B} |\Psi_n \rangle) = a_n (\hat{B} |\Psi_n \rangle) = a_n \hat{B} |\Psi_n\rangle$
 $\hat{A} (\hat{B} |\Psi_n \rangle) = a_n (\hat{B} |\Psi_n \rangle) = a_n \hat{B} |\Psi_n\rangle$
 $\hat{B} |\Psi_n \rangle = a_n |\Psi_n\rangle$
 $\hat{B} |\Psi_n \rangle = b_n |\Psi_n\rangle$
 $\hat{A} \hat{B} |\Psi_n \rangle = b_n |\Psi_n\rangle = \hat{B} \hat{A} |\Psi_n\rangle$

Fifth postulate

If the measurement of the physical quantity A gives the result a_n, the wavefunction of the system immediately after the measurement is the eigenvector

 \$\mathcal{M}_n\$

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Position and probability



particle in a one-dimensional hard box according to classical mechanics removed for copyright reasons. See Mortimer, R. G. *Physical Chemistry*. 2nd ed. San Diego, CA: Elsevier, 2000, page 555, Figure 15.3.

Graphs of the probability density for positions of a

Diagram showing the probability densities of the first 3 energy states in a 1D quantum well of width L.

"Collapse" of the wavefunction



Figure by MIT OpenCourseWare.

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Quantum double-slit

Image removed due to copyright restrictions. Please see any experimental verification of the double-slit experiment, such as http://commons.wikimedia.org/wiki/Image:Doubleslitexperiment_results_Tanamura_1.gif

Image of a double-slit experiment simulation removed due to copyright restrictions. Please see "Double Slit Experiment." in *Visual Quantum Mechanics.*

Deterministic vs. stochastic

- Classical, macroscopic objects: we have welldefined values for all dynamical variables at every instant (position, momentum, kinetic energy...)
- Quantum objects: we have well-defined probabilities of measuring a certain value for a dynamical variable, when a large number of identical, independent, identically prepared physical systems are subject to a measurement.

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When scientists turn bad...



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Cat wavefunction

$$|\Psi_{cat}(t)\rangle = |\Psi_{alive}\rangle \left(\exp\left(-\frac{t}{\tau}\right)\right)^{\frac{1}{2}} + |\Psi_{dead}\rangle \left(1 - \exp\left(-\frac{t}{\tau}\right)\right)^{\frac{1}{2}}$$

 There is not a value of the observable until it's measured (a conceptually different "statistics" from thermodynamics)

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Uncertainties, and Heisenberg's Indetermination Principle $(A \ge : \int \gamma^{*} (A \land \gamma) d\vec{r}$ $(\Delta A)^{2} = \langle (A - \langle A \rangle)^{2} \rangle = \langle A^{2} \rangle - \langle A \rangle^{2}$ $\Delta A \Delta B \ge \frac{1}{2} |\langle [A, B] \rangle|$ $\left[x, -i\hbar \frac{d}{dx} \right] = i\hbar$

Linewidth Broadening

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Please see: Fig. 2 in Uhlenberg, G., et al. "Magneto-optical Trapping of Silver Ions." *Physical Review A* 62 (November 2000): 063404.

 $\Delta E \Delta t \ge \frac{\hbar}{2}$

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Top Three List

- Albert Einstein: "Gott wurfelt nicht!" [God does not play dice!]
- Werner Heisenberg "I myself . . . only came to believe in the uncertainty relations after many pangs of conscience. . ."
- Erwin Schrödinger: "Had I known that we were not going to get rid of this damned quantum jumping, I never would have involved myself in this business!"



Figure by MIT OpenCourseWare.

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Commutation Relation

$$\hat{L} \quad \hat{L}_{x} \quad \hat{L}_{y} \quad \hat{L}_{z}$$

$$\begin{bmatrix} \hat{L}^{2}, \hat{L}_{x} \end{bmatrix} = \begin{bmatrix} \hat{L}^{2}, \hat{L}_{y} \end{bmatrix} = \begin{bmatrix} \hat{L}^{2}, \hat{L}_{z} \end{bmatrix} = 0$$

$$\begin{bmatrix} \hat{L}_{x}, \hat{L}_{y} \end{bmatrix} = i\hbar\hat{L}_{z} \neq 0$$

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Angular Momentum in Spherical Coordinates

$$\hat{L}_{z} = -i\hbar \frac{\partial}{\partial \varphi}$$
$$\hat{L}^{2} = -\hbar^{2} \left(\frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{\sin^{2} \theta} \frac{\partial^{2}}{\partial \varphi^{2}} \right)$$





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Same as a beating drum...



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... for the career helioseismologist



Image courtesy of NSO/AURA/NSF. Used with permission. Normal modes (i.e. sound, or seismic waves) for the Sun (basically jello in a 3d spherical box) 3.23 Electronic, Optical and Magnetic Properties of Materials - Nicola Marzari (MIT, Fall 2007)

Angular Momentum, then...



Figure by MIT OpenCourseWare.

$$L^2 = l(l+1)\hbar^2 = 0, \quad 2\hbar^2, \quad 6\hbar^2...$$

 $L_z = 0, \quad \pm \hbar, \quad \pm 2\hbar, \quad \pm 3\hbar...$

An electron in a central potential (I)



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An electron in a central potential (III)

$$u_{nl}(r) = r R_{nl}(r) \qquad V_{eff}(r) = \frac{\hbar^2}{2\mu} \frac{l(l+1)}{r^2} - \frac{Ze^2}{4\pi\varepsilon_0 r}$$
$$\left[-\frac{\hbar^2}{2\mu} \frac{d^2}{dr^2} + V_{eff}(r) \right] u_{nl}(r) = E_{nl} u_{nl}(r)$$

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What is the $V_{eff}(r)$ potential ?



Figure by MIT OpenCourseWare.





Figures by MIT OpenCourseWare.

 \mathbf{R}_{10} $r^2 R^2_{\ 10}$ 2 0.4 1 0.2 0 0 2 2 3 3 0 4 0 R₂₀ $r^2 R^2_{\ 20}$ 0.6 0.15 0.4 0.1 0.2 10 0.05 0 -0.2 0 r 10 $r^2 R^2_{\ 21}$ R_{21} 0.12 0.15 0.08 0.1 0.04 0.05 0 0 6 10 r 0 2 6 10 0 R₃₀ $r^2 R^2_{\ 30}$ 0.4 0.08 0.2 8 12 0.04 16 0 -0.1 0 12 16 R_{31} $r^2 R^2_{31}$ 0.08 0.8 0.04 0.4 0 -0.04 0 16 r 0 8 12 4 $r^2 R^2_{\ 32}$ R_{32} 0.04 0.8 0.02 0.4 0 0 16 r 12 16 12 0 4 8 0 4 8 Radial functions $R_{nl}(r)$ and radial distribution functions $r^2 R^2_{nl}(r)$ for

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Radial functions $R_{nl}(r)$ and radial distribution functions $r^2R^2_{nl}(r)$ for atomic hydrogen. The unit of length is $a_{\mu} = (m/_{\mu}) a_0$, where a_0 is the first Bohr radius.

The Grand Table

Shell	Quant n	um n l	umbers <i>m</i>	Spectroscopic notation	Wave function $\Psi_{nlm}(\mathbf{r}, \mathbf{\theta}, \mathbf{\phi})$
K	1	0	0	1s	$\frac{1}{\sqrt{\pi}} (Z/a_0)^{3/2} \exp(-Zr/a_0)$
L	2	0	0	2s	$\frac{1}{2\sqrt{2\pi}}(Z/a_0)^{3/2}(1-Zr/2a_0)\exp(-Zr/a_0)$
	2	1	0	2p ₀	$\frac{1}{4\sqrt{2\pi}}(Z/a_0)^{3/2}(Zr/a_0)\exp(-Zr/2a_0)\cos\theta$
	2	1	± 0	$2p_{\pm 1}$	$\mp \frac{1}{8\sqrt{\pi}} (Z/a_0)^{3/2} (Zr/a_0) \exp(-Zr/2a_0) \sin\theta \exp(\pm i\phi)$
М	3	0	0	3s	$\frac{1}{3\sqrt{3\pi}}(Z/a_0)^{3/2}(1-2Zr/3a_0+2Z^2r^2/27a)\exp(-Zr/3a_0)$
	3	1	0	3p ₀	$\frac{2\sqrt{2}}{27\sqrt{\pi}}(Z/a_0)^{3/2}(1-Zr/6a_0)(Zr/a_0)\exp(-Zr/3a_0)\cos\theta$
	3	1	± 1	3p _{±1}	$\mp \frac{2}{27\sqrt{\pi}} (Z/a_0)^{3/2} (1-Zr/6a_0) (Zr/a_0) \exp(-Zr/3a_0) \sin \theta \exp(\pm i\phi)$
	3	2	0	3d ₀	$\frac{1}{81\sqrt{6\pi}}(Z/a_0)^{3/2}(Z^2r^2/a) \exp(-Zr/3a_0) (3\cos^2\theta - 1)$
	3	2	± 1	$3d_{\pm 1}$	$\overline{+}_{81\sqrt{\pi}}^{1}(Z/a_0)^{3/2}(Z^2r^2/a) \exp(-Zr/3a_0)\sin\theta\cos\theta\exp(\pm i\phi)$
	3	2	± 2	3d _{±2}	$162\sqrt{\pi} (Z/a_0)^{3/2} (Z^2 r^2/a) \exp(-Zr/3a_0) \sin^2\theta \exp(\pm 2i\phi)$

The complete normalised hydrogenic wave functions corresponding to the first three shells, for an 'infinitely heavy' nucleus. The quantity $a_0 = 4\pi\epsilon_0 h^2/me^2$ is the first Bohr radius. In order to take into account the reduced mass effect one should replace a_0 by $a_u = a_0 (m/\mu)$

Solutions in the central Coulomb Potential: the Alphabet Soup

http://www.orbitals.com/orb/orbtable.htm



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